



PROGRESS



PRediction Of Geospace Radiation Environment and Solar wind parametersS

New EC Horizon 2020 funded project.

PARTICIPANTS



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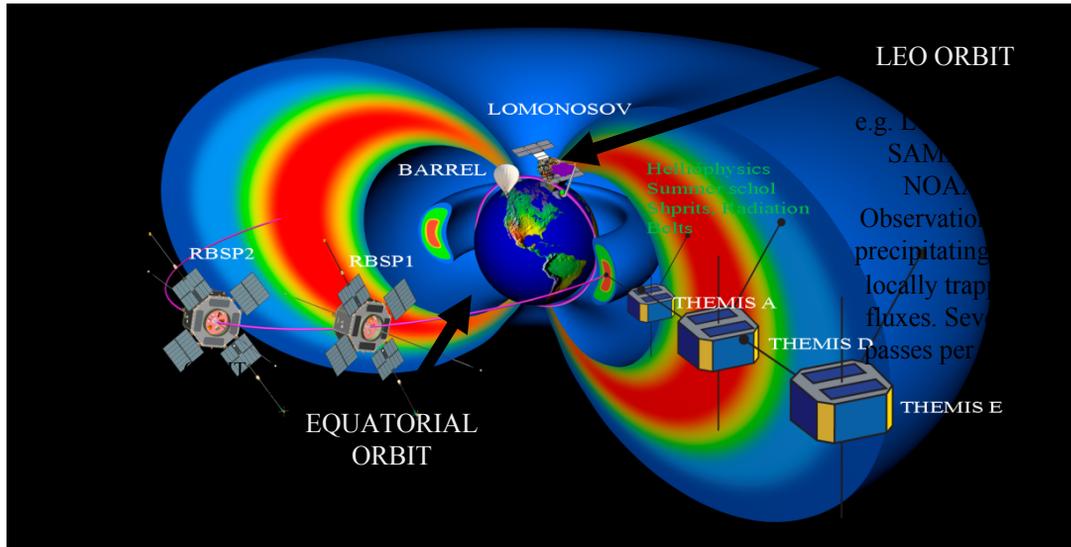
IRF-L



AIMS

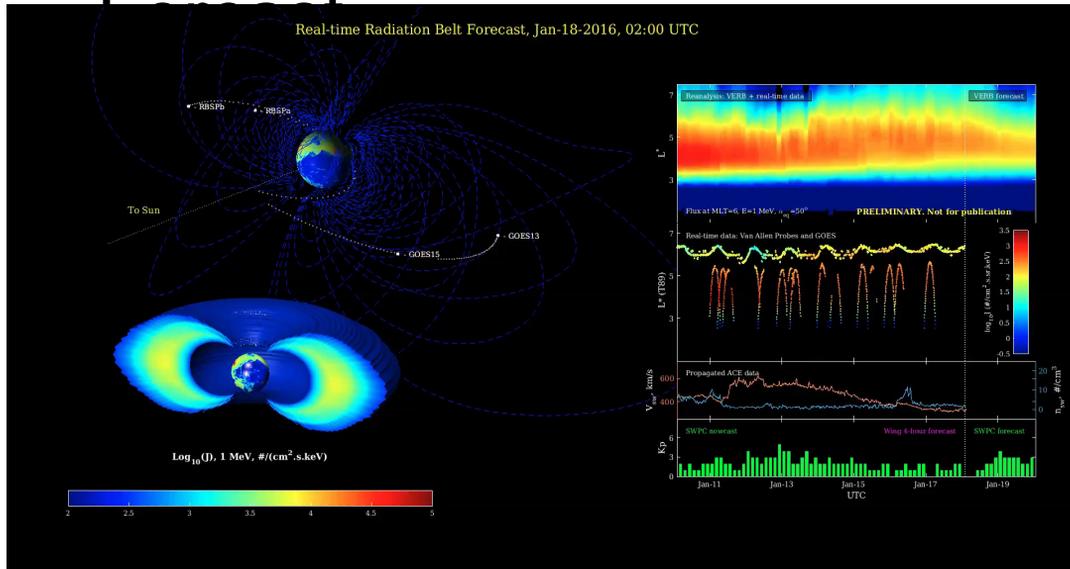
- Development of a European Solar Wind model
- Models for the evolution of geomagnetic indices
- Statistical Wave models of wave activity
- Development and coupling of systems methodologies with physically based models
- Tools for robust, reliable forecasts for
 - geomagnetic indices
 - particle environment of the inner magnetosphere

Multi-Point Observations



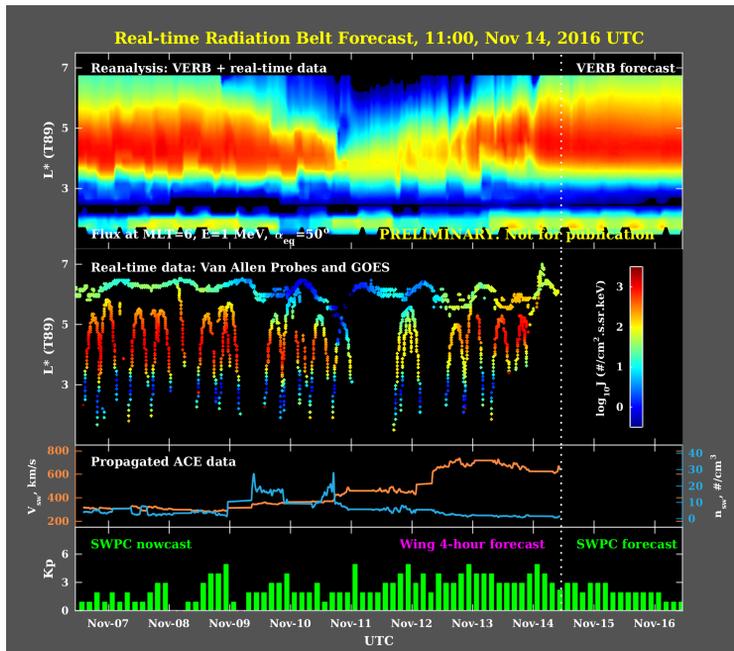
Radiation Belts Assimilative

Data from different satellites is blended with a physics-based model

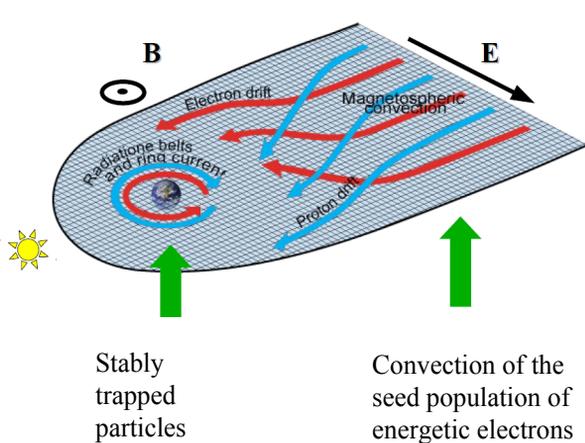


Radiation Belts Assimilative Forecast

<http://rbm.epss.ucla.edu/realtime-forecast/>



Particle Trajectories of Ring Current and Radiation Belt Particles



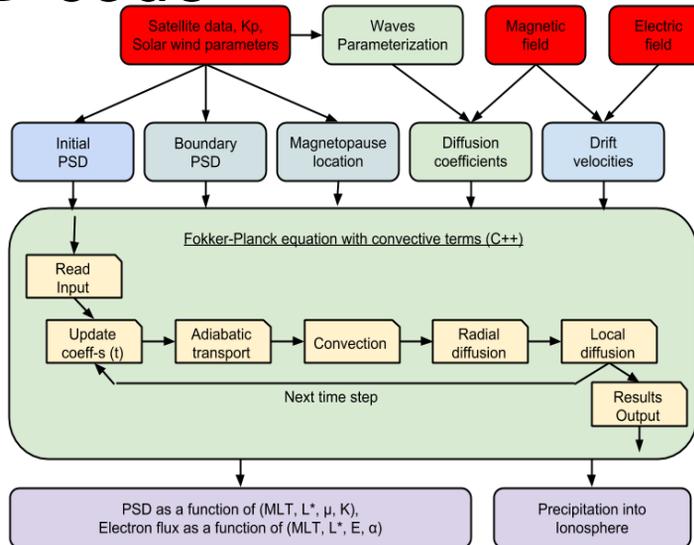
Drift of lower energy particles is dominated by $E \times B$ drift.

Radiation Belt particles are subject to the gradient and curvature drifts and will drift around the Earth.

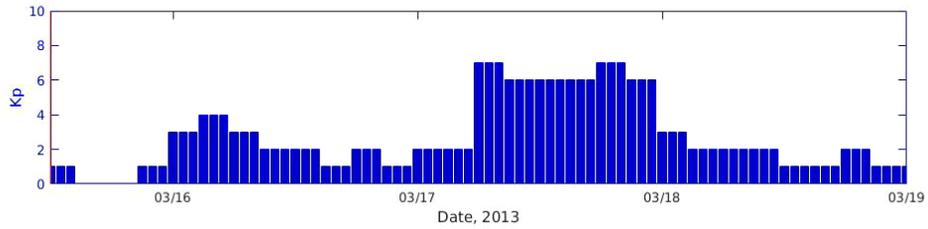
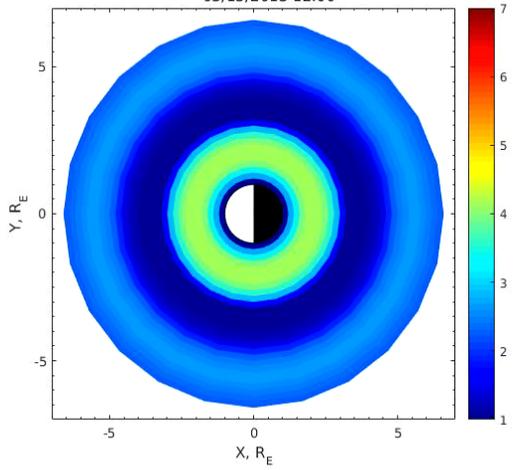
Electrons – eastward,

Ions – westward.

Block Diagram Showing Data Exchange between the modules of the VERB 4D code



Flux, $\log_{10}(\#/s/cm^2/sr/keV)$, $E = 0.2$ MeV, $\alpha = 50$ deg
03/15/2013 12:00



Comparison with IMAGE EUV

26-Jun-2001 22:27:00 UT

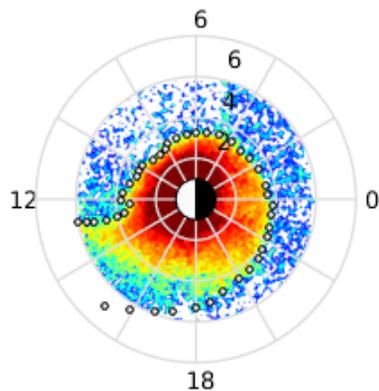
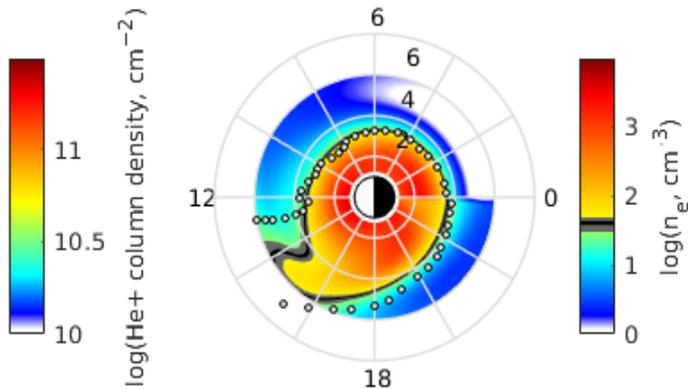
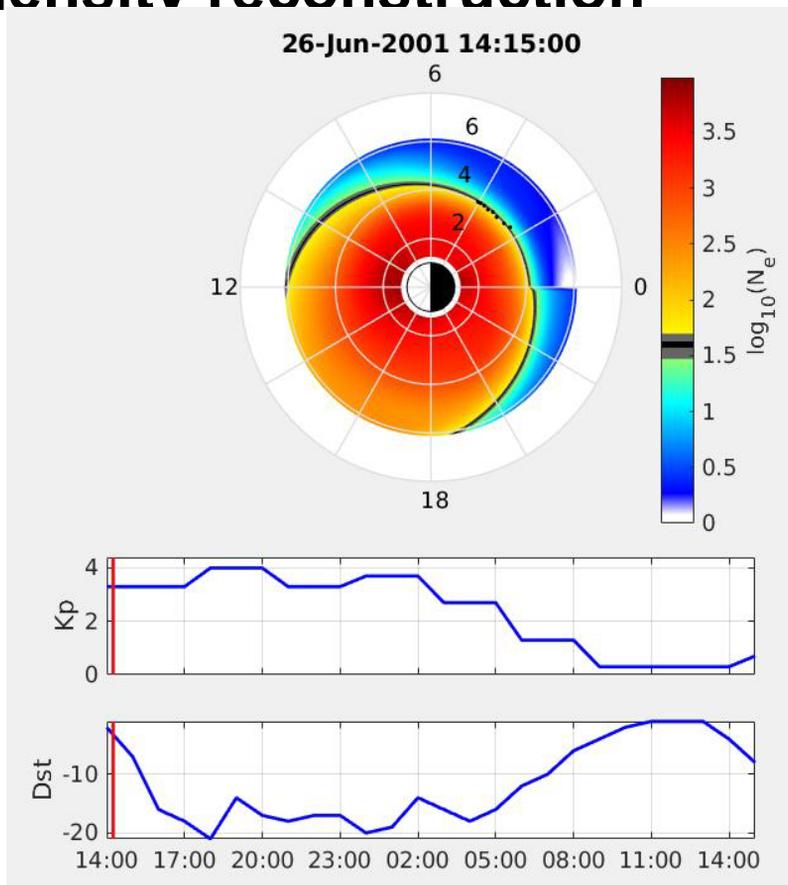


IMAGE EUV



Neural network output

Global density reconstruction



Questions:

Why are we in such a big room?

What do you think we should be doing?

What are other products that may be useful?

Should we focus on predictions or on reanalysis?

How our predictions may help?

VERB-4D code development

Shprits Y., Kellerman A., Drozdov A., Aseev N.

VERB-4D code testing

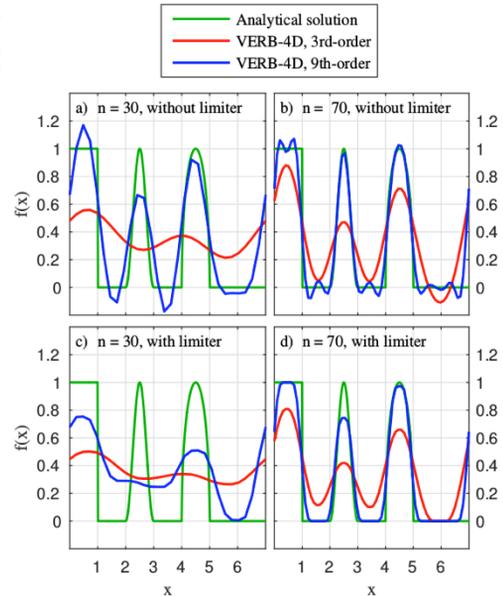
The VERB-4D code was thoroughly validated [Aseev et al., 2016].

Implemented numerical schemes are stable regardless time step and computational grid.

Solutions of advection and diffusion equations converge to analytical solutions.

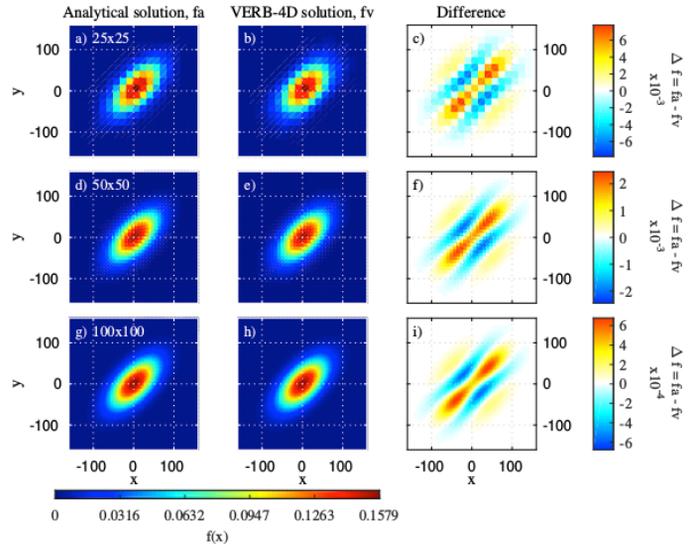
VERB-4D code testing

Implemented 9th-order scheme shows much more accurate results than the 3rd-order one.



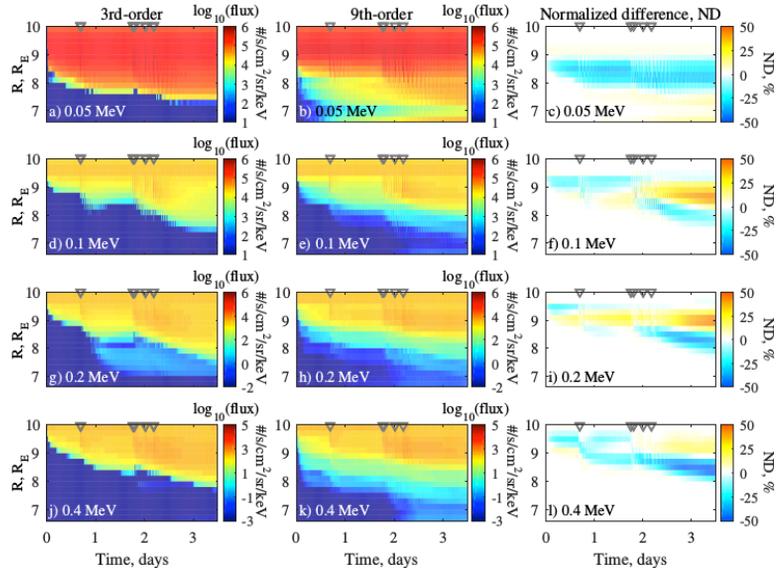
VERB-4D code testing

The solution of diffusion equation with cross-derivative terms converges to the analytical one.



VERB-4D code testing

The order of numerical scheme may significantly affect the results of charged particles advection modeling



Boundary conditions for extended VERB-4D code

VERB-4D code is now being extended up to radial distances $R \sim 10-15 R_e$.

Accurate boundary conditions are required in the inner central plasma sheet

$$f = N \left(\frac{M}{2\pi\kappa E_0} \right)^{3/2} \frac{\Gamma(\kappa+1)}{\Gamma(\kappa-1/2)} \left[1 + \frac{E^*}{\kappa E_0} \right]^{-\kappa-1}$$

Omnidirectional distribution function with kappa

$$E_0 = kT \frac{\kappa-3/2}{\kappa}$$

Boundary condition parameterization

Plasma sheet electron density and temperature are taken from [Dubya et al., 2016]

$$N_{ps} = A_1 + A_2 R^* + A_3 \phi^{*2} R^* + A_4 \phi^{*2} + A_5 N_{sw}^* + (A_6 + A_7 R^*) B_S^*$$

$$T_{ps} = [A_1 + A_2 \phi^* + A_3 V_{sw}^* + (A_4 + A_5 \phi^{*2} R^*) B_S^{*A_7} + A_6 R^* B_N^{*A_8}]^{A_9}$$

$$\phi^* = \phi/90^\circ \quad \phi = \arctan(-Y_{GSM}/X_{GSM}) \quad R^* = R/10R_E \quad N_{sw}^* \quad V_{sw}^* \quad B_S^* \quad B_N^*$$

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averaged over preceding time interval (specific for each solar wind parameter).

Boundary condition parameterization

$K = 4.1$ is determined by minimization of RMSE between THEMIS fluxes and kappa distribution function

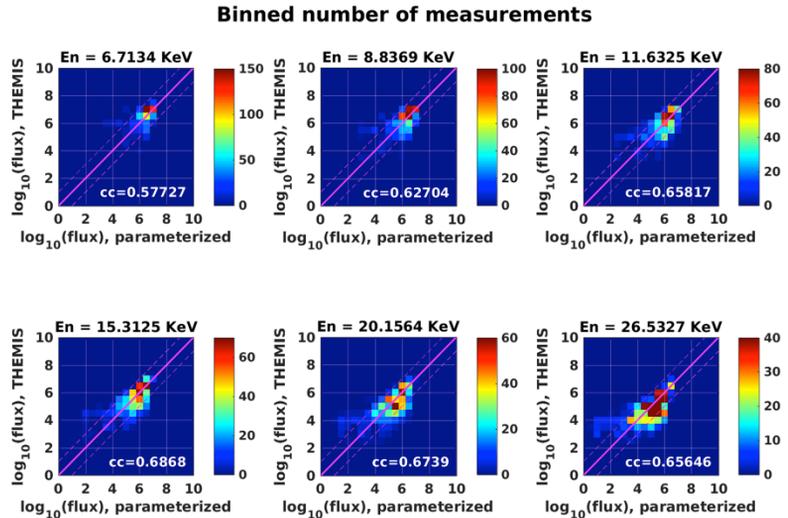


Figure. Predicted plasma sheet electron fluxes versus THEMIS ESA measurements

Boundary condition parameterization

Normalized occurrence

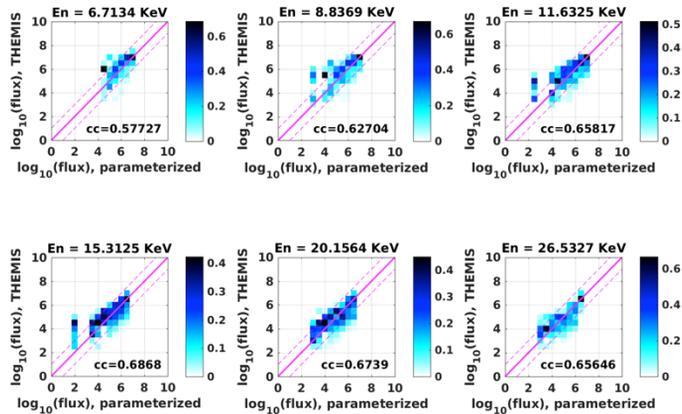


Figure. Conditional probability of predicted flux to be in particular bin of THEMIS data.